

Potential Field Source Surface Model and Solar Wind Prediction Hansen Experimental Physics Laboratory, Stanford University



Solar wind Prediction

Arge and Pízzo (2000)

 $V_{sw} = 267.5 + 410.0/f^{(1.0/2.5)}$ (3) f => FTE calculated using PFSS model

Prediction not always agree with observation

Potential Field Source Surface Model

Schatten, Wilcox and Ness 1969 Altschuler and Newkirk 1969

assumptions

- little current flows between photosphere and source surface
- coronal magnétic field can be derived from a potential obeying LaPlace's equation
- at the source surface all field lines are radial



PFSS Model: Parameters

height of source surface radius of the inner sphere number of multipole components in spherical harmonic expansion: Nmax

2.5 Rsun

· 1.0 Rsun

N_{max} = 9, ... 90 WSO data: 22 Kítt Peak: 90

PFSS Model: Limitations

very sensitive to rapid field evolutions

- magnetic field predicted for mid- and high latitudes does not agree with observations potential field approximation not strictly valid for solar corona
- exclusion of current sheet causes severe discrepancies
- location of source surface, 2.5 R_{sun}, much lower than Alfvén crítical point





Levíne, Altschuler and Harvey (1977)

inverse correlation between SWS observed at 1 AU and FTE on the source surface, using PFSS model





Flux Tube Expansion Factor (FTE) $FTE = R_s/R_{ss} * B_r(\theta_{ss}, \Phi_{ss})/B_r(\theta_s, \Phi_s)$ (1)

B_r(θ_s,Φ_s) photospheric magnetic field
R_s photospheric radius
B_r(θ_{ss},Φ_{ss}) source surface magnetic field

Rss source surface radius



Wang and Sheeley (1990; 1994; 1997) Confirmed the inverse correlation

Speed	FTE
< 450	> 20
450-550	10-20
550-650	8-10
650-750	4.5-8
> 750	< 4.5



Inverse Mapping

Computation of correlation between FTE and SWS involves:

1. determination of precise location of coronal sources of solar wind $\Theta_0 = \Theta_R$ $\Phi_0 = \Phi_R + \omega R_E / V_R$ (2)

2. Identification of photospheric footpoints of these sources by tracing along the magnetic field lines



Discrepancy - Causes

- quality and resolution of photospheric data
- existence of transients not included in the model
- stream-stream interaction
- inverse mapping of solar wind to its source
 - límitations of PFSS model itself

How to improve?

Arge and Pízzo (2000)

improved photospheric field data and allowed streamstream interaction

discrepancies still exist

What Next?

Look at calculation of FTE

we investigated the sources of errors caused in FTE computation using PFSS model

Current Sheet Source Surface Model Improvements over SS model cusp surface: field lines are open but NOT necessarily radial; includes effects of streamer current sheets source surface: placed near the Alfvén critical point

uses source surface technique to include effects of volume current beyond source surface







FTE and Mmax



Variation of FTE with Nmax used in PFSS model







Optimm Mmax





Inverse Mapping - Correlation Coefficient



variation of correlation coefficient with dífferent values of solar wind transit time adopted in the inverse mapping technique

RD IIN 1891

Scatter Plot - FTE vs SWS





Histogram: "Error" in Longitude

+00

Ο.

Ο

1978:1981

1988:1991

2C

All data

All data

Shift in longitude (deg)

-30



